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Goldstone Bosons and Global Strings in a Warped Resolved Conifold

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ABSTRACT

A warped resolved conifold background of type IIB theory, constructed in hep-th/0701064, is dual to the supersymmetric $SU(N) \times SU(N)$ gauge theory with a vacuum expectation value (VEV) for one of the bifundamental chiral superfields. This VEV breaks both the superconformal invariance and the baryonic symmetry. The absolute value of the VEV controls the resolution parameter of the conifold. In this paper we study the phase of the VEV, which corresponds to the Goldstone boson of the broken symmetry. We explicitly construct the linearized perturbation of the 4-form R-R potential that contains the Goldstone boson. On general grounds, the theory should contain global strings which create a monodromy of the pseudoscalar Goldstone boson field. We identify these strings with the $D3$ -branes wrapping the two-cycle at the tip of the warped resolved conifold.

1 Introduction and summary

The AdS/CFT duality [1, 2, 3] has produced major progress in our understanding of the intimate relationship between the dynamics of gauge theories and strings. The basic version of the duality is motivated by considering a stack of many $D3$ -branes in flat space. The dual descriptions of $D3$ -branes as a supergravity background solution on which strings propagate, or as objects carrying a worldvolume gauge theory, allows one to identify type IIB strings on $AdS_5 \times S^5$ with $\mathcal{N} = 4$ SYM theory. Furthermore one can engineer more elaborate (and less symmetric) versions of the duality by considering $D3$ -branes at the tip of a generic CY cone. Among the simplest examples is the cone over $T^{1,1}$, the conifold, which leads to a duality between type IIB strings on $AdS_5 \times T^{1,1}$ with N units of R-R 5-form flux and a $SU(N) \times SU(N)$ SCFT coupled to bifundamental chiral superfields A_1, A_2, B_1, B_2 [4].

The singularity of the conifold can be smoothed out in two different ways: either by blowing up a 3-cycle, leading to the deformed conifold, or by blowing up a 2-cycle, leading to the resolved conifold. The deformation of the conifold may be produced by turning on M units of R-R flux through the 3-cycle [5, 6] (for reviews, see [7]). The resulting warped product of $R^{3,1}$ and the deformed conifold gives rise to interesting phenomena arising from the fact that the corresponding $SU(N+M) \times SU(N)$ gauge theory undergoes a duality cascade leading to confinement and chiral symmetry breaking in the infrared. The resolution of the conifold has a simpler interpretation: it corresponds to giving expectation values to the fields A_i, B_j in the $SU(N) \times SU(N)$ gauge theory [8]. Recently a particularly simple example where only one of the fields acquires a diagonal VEV, $B_2 = uI_{N \times N}$, was worked out in detail [9]. The dual warped product of $R^{3,1}$ and the resolved conifold corresponds to a stack of N $D3$ -branes placed at a point on the blown-up 2-sphere at the tip. The warp factor produced by the $D3$ -branes was explicitly solved for.

The VEV of B_2 spontaneously breaks the $U(1)_B$ symmetry of the theory. The gauge invariant order parameter for this breaking is the baryonic operator $\det B_2$. In [9] the modulus of the VEV of this operator was computed holographically using a Euclidean $D3$ -brane wrapping a holomorphic 4-chain. In this paper we will discuss the phase of this VEV and phenomena associated with its variation. One of these phenomena is the presence of a Goldstone boson; we will exhibit the normalizable 4-form perturbation around the warped resolved conifold that contains it. The required perturbation contains a product of two 2-forms, an antisymmetric tensor in $R^{3,1}$ and a closed form W on the warped resolved conifold. The dual of the antisymmetric tensor defines the pseudoscalar Goldstone boson p . We will write down the equations that determine the 2-form W and show that they come from the minimization of a positive definite functional which, in the warped case and with appropriate boundary conditions, remains finite. In fact, this functional is the norm of the form W . The asymptotically AdS warp factor is crucial for the normalizability of W ; in the unwarped case, the corresponding W is not normalizable.

On general grounds, a broken global symmetry may give rise to “global” strings associated with a non-trivial monodromy of the Goldstone boson. We will show that, on the string side of the duality, they are realized as $D3$ -branes which wrap the finite S^2 at the tip of the warped resolved conifold, with the remaining two world volume directions lying within $R^{3,1}$. Interestingly the tension of these strings is not sensitive to the warp factor. We will show that they are BPS saturated and are stable at the tip, $r = 0$. Such wrapped $D3$ -branes couple to the components of R-R 4-form described above, and create the monodromy of the pseudoscalar p . We will also show that the phase of the baryonic condensate computed holographically from $e^{-S_{E3}}$ is determined by p .

The outline of the paper is as follows. In section 2 we review the duality between the warped resolved conifold and the gauge theory which was elucidated in [9]. In section 3 we study global strings obtained from partially wrapped $D3$ -branes and check their κ -symmetry. In section 4 we derive the equations for the fluctuation of the R-R 4-form potential containing the Goldstone mode of the broken $U(1)_B$ symmetry, and solve them in various limits. Section 5 is devoted to a further analysis of spontaneous breaking of

$U(1)_B$ using the dual supergravity. In particular, we calculate the Goldstone boson ‘‘decay constant.’’ We conclude with a few comments in section 6.

2 The Warped Resolved Conifold

It is well-known that the conifold can be described by the equation

$$z_1^2 + z_2^2 + z_3^2 + z_4^2 = 0 , \quad z_i \in \mathbb{C} . \quad (1)$$

By defining a related set of w_i coordinates, we can write this equation as

$$Z = \begin{pmatrix} z_3 + iz_4 & z_1 - iz_4 \\ z_1 + iz_4 & -z_2 + iz_4 \end{pmatrix} = \begin{pmatrix} w_1 & w_3 \\ w_4 & w_2 \end{pmatrix} ; \det Z = 0 . \quad (2)$$

It is possible to solve this by choosing

$$Z = \begin{pmatrix} a_1 b_1 & a_1 b_2 \\ a_2 b_1 & a_2 b_2 \end{pmatrix} \quad (3)$$

and use the (a_i, b_i) coordinates to describe the conifold. However they are not uniquely determined, since they are identified under

$$(a_i, b_i) \rightarrow (\lambda a_i, \frac{1}{\lambda} b_i) , \quad \lambda \in C^* . \quad (4)$$

In order to partially fix this freedom, one can impose the phase identification $a_i \sim e^{i\alpha} a_i$ and $b_i \sim e^{-i\alpha} b_i$. In the dual gauge theory this corresponds to the $U(1)_B$ symmetry which assigns opposite charges to a_i and b_i . We are still left with the modulus of the above transformation, which we can fix by demanding $|b_1|^2 + |b_2|^2 - |a_1|^2 - |a_2|^2 = 0$. In the $U(1)$ gauge theory this constraint arises from the D-term, but as emphasized in [8], in the $SU(N) \times SU(N)$ case the analogous constraint is absent.

A simple way to understand the resolution of the conifold is to deform the modulus constraint above into

$$|b_1|^2 + |b_2|^2 - |a_1|^2 - |a_2|^2 = u^2 , \quad (5)$$

where u is a real parameter which controls the resolution. The resolution corresponds to a blow up of the S^2 at the bottom of the conifold. In the dual gauge theory turning on u corresponds to a particular choice of vacuum. After promoting the a, b fields into the bifundamental chiral superfields of the dual gauge theory, we can define the operator \mathcal{U} as

$$\mathcal{U} = \frac{1}{N} \text{Tr}(B_1^\dagger B_1 + B_2^\dagger B_2 - A_1^\dagger A_1 - A_2^\dagger A_2) . \quad (6)$$

Thus, the warped singular conifolds correspond to gauge theory vacua where $\langle \mathcal{U} \rangle = 0$, while the warped resolved conifolds correspond to vacua where $\langle \mathcal{U} \rangle \neq 0$. In the latter case, some VEVs for the bi-fundamental fields A_i, B_j must be present. Since these fields are charged under the $U(1)_B$ symmetry, the warped resolved conifolds correspond to vacua where this symmetry is broken [8].

A particularly simple choice is to give a diagonal VEV to only one of the scalar fields, say, B_2 . As seen in [9], this choice breaks the $SU(2) \times SU(2) \times U(1)_B$ symmetry of the CFT down to $SU(2) \times U(1) \times U(1)$. The string dual is given by a warped resolved conifold

$$ds^2 = h^{-1/2} dx_{1,3}^3 + h^{1/2} ds_6^2 . \quad (7)$$

The explicit form of the Calabi-Yau metric of the resolved conifold is given by [10]

$$\begin{aligned} ds_6^2 = & K^{-1} dr^2 + \frac{1}{9} Kr^2 \left(d\psi^2 + \cos \theta_1 d\phi_1 + \cos \theta_2 d\phi_2 \right)^2 + \frac{1}{6} r^2 (d\theta_1^2 + \sin^2 \theta_1 d\phi_1^2) \\ & + \frac{1}{6} (r^2 + 6u^2) (d\theta_2^2 + \sin^2 \theta_2 d\phi_2^2), \end{aligned} \quad (8)$$

where

$$K = \frac{r^2 + 9u^2}{r^2 + 6u^2}. \quad (9)$$

The N $D3$ -branes sourcing the warp factor are located at the north pole of the finite S^2 , i.e. at $r = 0, \theta_2 = 0$. The corresponding warp factor h is a function of both r and θ_2 , which can be written as [9]

$$h = L^4 \sum_{l=0}^{\infty} (2l+1) H_l(r) P_l(\cos \theta_2) \quad (10)$$

where $L^4 = \frac{27\pi}{4} g_s N (\alpha')^2$, $P_l(\cos \theta)$ is the l -th Legendre polynomial, and

$$H_l = \frac{2C_\beta}{9u^2 r^{2+2\beta}} {}_2F_1\left(\beta, 1+\beta, 1+2\beta; -\frac{9u^2}{r^2}\right), \quad (11)$$

with the coefficients C_β and β given by

$$C_\beta = \frac{(3u)^{2\beta} \Gamma(1+\beta)^2}{\Gamma(1+2\beta)}, \quad \beta = \sqrt{1 + \frac{3}{2} l(l+1)}. \quad (12)$$

Using the Euclidean $D3$ -brane located at fixed θ_2, ϕ_2 , it is possible to compute the VEV of the baryonic operator $\det B_2$. Its modulus is found to be $\sim u^{\frac{3N}{4}}$ [9]. Its phase, which is determined by value of the Goldstone boson of the broken $U(1)_B$ symmetry, will be discussed in section 5.2.

Far in the IR the gauge theory flows to the $\mathcal{N} = 4$ SYM theory, as evidenced by the appearance of an $AdS_5 \times S^5$ throat near the location of the stack of the $D3$ -branes. We will see that the gauge theory also contains an interesting additional sector coupled to this infrared CFT. The coupling of such an extra sector and an infrared CFT is reminiscent of the unparticle physics scenarios [11].

3 Global Strings

In addition to the existence of a Goldstone boson, a hallmark of a broken $U(1)$ symmetry is the appearance of “global” strings, around which the Goldstone boson carries a non-trivial monodromy. We will show that on the supergravity side of the duality these global strings are partially wrapped $D3$ -branes.

Let us consider the IR (small r) region. There the warp factor will approach some function g^2 whilst the resolved conifold contains a 2-sphere parametrized by θ_2, ϕ_2 :

$$ds^2 \rightarrow g^{-1} dx_{1,3}^2 + gu^2 d\Omega_2. \quad (13)$$

Let us now consider a $D3$ -brane whose world-volume spans the t, x, Ω_2 coordinates, such that the brane wraps the two-sphere but remains extended along one of the x directions. From the point of view of the dual gauge theory this will correspond to a string-like object along t, x . It is natural to identify it with the string originating from the breaking of $U(1)_B$, since its existence is connected with the finite S^2 at the bottom of the resolved conifold, which in turn requires that the $U(1)_B$ is broken.

It is straightforward to compute the tension of such a wrapped $D3$ -brane¹

$$T_s = 4\pi T_3 u^2 = \frac{u^2}{2\pi^2 g_s} . \quad (14)$$

Interestingly, the tension is completely independent of the warp factor and remains finite at the bottom of the conifold.² The global strings and the Goldstone modes they couple to belong to another sector of the theory which “sees” the whole S^2 , and is coupled to the $\mathcal{N} = 4$ SYM . The very existence of this extra sector coupled to the CFT is an interesting fact. It is reminiscent of the unparticle physics scenarios [11] typically characterized by a particle sector coupled to a conformal (unparticle) sector. Our construction amounts to a UV completion of such an unparticle scenario in terms of the $SU(N) \times SU(N)$ SCFT with a VEV for one of the bi-fundamental fields.

3.1 Kappa-symmetry for the $D3$

The kappa-symmetry projection which our $D3$ -brane should satisfy is

$$\Gamma_\kappa \epsilon = i\mathcal{L}_{DBI}^{-1} \gamma_4 \epsilon , \quad (15)$$

where γ_4 is the pull-back of the target space gamma matrices to the worldvolume of the brane, and ϵ is the background Killing spinor. Those spinors can be written in terms of the Killing spinor in the resolved conifold. For that, write the resolved conifold metric in terms of

$$\begin{aligned} e^1 &= \frac{r}{\sqrt{6}} d\theta_1 , & e^3 &= \frac{\sqrt{r^2+6u^2}}{\sqrt{6}} d\theta_2 , \\ e^2 &= \frac{r}{\sqrt{6}} \sin \theta_1 d\phi_1 , & e^4 &= \frac{\sqrt{r^2+6u^2}}{\sqrt{6}} \sin \theta_2 d\phi_2 , \\ e^5 &= \frac{\sqrt{K}r}{3} (d\psi + \cos \theta_1 d\phi_1 + \cos \theta_2 d\phi_2) , & e^0 &= K^{-\frac{1}{2}} dr ; \end{aligned}$$

so that the 6d metric reduces to

$$ds_6^2 = (e^0)^2 + (e^1)^2 + (e^2)^2 + (e^3)^2 + (e^4)^2 + (e^5)^2 , \quad (16)$$

It is then straightforward to see that for our brane wrapping t, x, θ_2, ϕ_2 and sitting at $r = 0$, the kappa-symmetry projector above reduces to

$$\Gamma_\kappa \epsilon = i\Gamma_{tx} \Gamma_{34} \epsilon , \quad (17)$$

where Γ_i are the flat space gamma matrices.

The background Killing spinor satisfies $\Gamma_{34} \epsilon = \Gamma_{12} \epsilon$ [12] (for a review see [13]). Since $i\Gamma_{34}$ and $i\Gamma_{12}$ are both projectors (they square to the identity and are traceless), these two equations can be solved to give

$$i\Gamma_{34} \epsilon = \epsilon , \quad i\Gamma_{12} \epsilon = \epsilon ,$$

so in the case at hand the kappa-symmetry projection reduces to

$$\Gamma_\kappa \epsilon = \Gamma_{tx} \epsilon = \epsilon . \quad (18)$$

Now since Γ_{tx} is a traceless matrix which squares to one, this condition is satisfied for half of the background spinors, so our strings are one-half BPS.

¹ We set $\alpha' = 1$ throughout the paper.

² In section 5 we will show that, due to backreaction, the tension of the string receives an additional logarithmically divergent contribution. This divergence is typical for an isolated co-dimension 2 global string.

4 The Fluctuation Containing the Goldstone Mode

In order to gain more understanding of the strings we have found, we can consider the linearized back-reaction in the background caused by our probe $D3$. Such a brane will source, to linearized order, a fluctuation in the 4-form RR potential containing the term $a_2(x) \wedge W$ where a_2 is a 2-form in $R^{3,1}$ and W is a closed 2-form in the resolved conifold: $dW = 0$. This perturbation has to satisfy the linearized equations of motion, which read $d\delta F^{(5)} = 0$ and $\delta F^{(5)} = \star\delta F^{(5)}$. We can ensure the latter by taking

$$\delta F^{(5)} = (1 + \star)d(a_2(x) \wedge W). \quad (19)$$

Then the equations of motion reduce to

$$d \star_4 da_2 = 0, \quad (20)$$

provided W satisfies

$$d(h \star_6 W) = 0, \quad (21)$$

where \star_4, \star_6 are the Hodge duals with respect to the unwarped Minkowski and resolved conifold metrics, respectively. Introducing the field $p(x)$ through $\star_4 da_2 = dp$, we note that the fluctuation in the 5-form field strength reads

$$\delta F^{(5)} = da_2 \wedge W + dp \wedge h \star_6 W. \quad (22)$$

The corresponding fluctuation of the 4-form potential is

$$\delta C^{(4)} = a_2(x) \wedge W + p h \star_6 W. \quad (23)$$

We will see later on that the field $p(x)$ is the Goldstone boson for the broken $U(1)_B$.

4.1 Equations for W

Different types of forms in conifolds have been discussed in the literature (see for example [12], [14], [15]). In our case, we are searching for a closed 2-form which is co-closed upon multiplication with the warp factor. We will consider an ansatz that satisfies $dW = 0$:

$$W = \sin \theta_2 d\theta_2 \wedge d\phi_2 + d(f_1 g^5 + f_2 \sin \theta_2 d\varphi_2), \quad (24)$$

where f_1, f_2 are functions of r, θ_2 . It is convenient to define a re-scaled set of vielbeins

$$\begin{aligned} \epsilon_1^1 &= d\theta_1, & \epsilon_2^1 &= d\theta_2, \\ \epsilon_1^2 &= \sin \theta_1 d\phi_1, & \epsilon_2^2 &= \sin \theta_2 d\phi_2, \end{aligned}$$

and

$$g^5 = d\psi + \cos \theta_1 d\phi_1 + \cos \theta_2 d\phi_2. \quad (25)$$

Then the 6-dimensional metric on the resolved conifold becomes

$$ds_6^2 = K^{-1} dr^2 + \frac{Kr^2}{9}(g^5)^2 + \frac{r^2}{6}(\epsilon_1^1)^2 + \frac{r^2}{6}(\epsilon_1^2)^2 + \frac{r^2 + 6u^2}{6}(\epsilon_2^1)^2 + \frac{r^2 + 6u^2}{6}(\epsilon_2^2)^2. \quad (26)$$

After some algebra, the condition $d(h \star_6 W) = 0$ translates into two coupled differential equations for f_1 and f_2 :

$$\begin{aligned} &\partial_r \left(\frac{h \partial_r f_1 r(r^2 + 6u^2)}{12} \right) + \frac{h}{3r(r^2 + 6u^2)} \left(r^4 (1 - f_1 + \frac{\partial_{\theta_2}(f_2 \sin \theta_2)}{\sin \theta_2}) - f_1(r^2 + 6u^2)^2 \right) \\ &+ \frac{r}{2K} \left(\frac{1}{\sin \theta_2} \partial_{\theta_2} [h \sin \theta_2 \partial_{\theta_2} f_1] \right) = 0; \end{aligned} \quad (27)$$

and

$$\partial_{\theta_2} \left[\frac{hr^3}{3(r^2 + 6u^2)} \left(1 - f_1 + \frac{\partial_{\theta_2}(f_2 \sin \theta_2)}{\sin \theta_2} \right) \right] + \partial_r \left[\frac{hKr^3}{18} \partial_r f_2 \right] = 0 . \quad (28)$$

These equations must be supplemented by a set of boundary conditions. Since the $\delta C^{(4)}$ fluctuation couples to the D3-brane wrapped over the S^2 at $r = 0$, W should approach there the volume form of the finite S^2 . For that we require both $f_1(r = 0, \theta_2), f_2(r = 0, \theta_2)$ to vanish. On the other hand, as we will see later on, in the large r region W should asymptote to the usual ω_2 2-form in the singular conifold, where $f_1(r \rightarrow \infty, \theta_2) = \frac{1}{2}$ while $f_2(r \rightarrow \infty, \theta_2) = 0$.

Interestingly equations (27)-(28) come from minimizing the following functional

$$I = \int W \wedge h \star_6 W = \int_0^\pi d\theta_2 \int_0^\infty dr \sin \theta_2 \left\{ \frac{hr^3}{3(r^2 + 6u^2)} \left(1 - f_1 + \frac{\partial_{\theta_2}(f_2 \sin \theta_2)}{\sin \theta_2} \right)^2 \right. \\ \left. + \frac{hr}{12} (r^2 + 6u^2) (\partial_r f_1)^2 + \frac{hr}{2K} (\partial_{\theta_2} f_1)^2 + \frac{hKr^3}{18} (\partial_r f_2)^2 + \frac{h}{3r} (r^2 + 6u^2) f_1^2 \right\} . \quad (29)$$

With the boundary conditions above, one can check that I remains finite using the warp factor in [9]. For small r the leading term in the warp factor, $h \sim \frac{1}{r^2} \delta(1 - \cos \theta_2)$, can be used to show that no divergence occurs. Had we assumed that f_1 was non-zero at the tip then I would not converge due to the last term. The same argument holds for f_2 , since otherwise I would diverge at $\theta_2 = \pi$.

For large r the leading behavior of the warp factor, $h = L^4/r^4$, renders the integral convergent provided f_1 approaches a constant (otherwise the term with $\partial_r f_1$ would blow up), while (27) sets it to 1/2 (in the unwarped case, the integral would instead diverge).

We have not been able to find an analytic solution for (27)-(28). However since with the chosen boundary conditions I remains finite, we expect that such a solution exists and is unique. It could be looked for using numerical relaxation algorithms applied to the functional I . We now turn to the analysis of (27)-(28) in several limits and check that they give sensible results.

4.2 Unwarped resolved conifold

If we set the warp factor to $h = 1$, then the equations we are solving describe the harmonic 2-form W_{harm} on the resolved conifold. This problem, in a slightly different context, has already been discussed in [15] with similar results. Note that in the case of $h = 1$, (29) is no longer convergent and therefore the solution for W is not normalizable.

In this instance we can take f_1, f_2 to depend only on r . Then (28) shows that f_2 may be set to zero, so we are left with a simplified version of (27) which reads

$$\partial_r \left(\frac{\partial_r f_1 r(r^2 + 6u^2)}{12} \right) + \frac{1}{3r(r^2 + 6u^2)} \left(r^4(1 - f_1) - f_1(r^2 + 6u^2)^2 \right) = 0 . \quad (30)$$

The solution to this equation is [15]

$$f_1 = \frac{r^2}{2(r^2 + 6u^2)} , \quad (31)$$

giving

$$W_{\text{harm}} = \left(\frac{1}{2} + \frac{3u^2}{r^2 + 6u^2} \right) \sin \theta_2 d\theta_2 \wedge d\phi_2 - \left(\frac{1}{2} - \frac{3u^2}{r^2 + 6u^2} \right) \sin \theta_1 d\theta_1 \wedge d\phi_1 + \frac{6u^2 r}{(r^2 + 6u^2)^2} dr \wedge g_5 . \quad (32)$$

In the UV (at large r), this form approaches the harmonic 2-form on the singular conifold,

$$\omega_2 = \frac{1}{2} (\sin \theta_2 d\theta_2 \wedge d\phi_2 - \sin \theta_1 d\theta_1 \wedge d\phi_1) . \quad (33)$$

Thus the harmonic form W_{harm} on the resolved conifold interpolates between $\sin \theta_2 d\theta_2 \wedge d\phi_2$, which is the volume form of the S^2 at $r = 0$, and ω_2 at large r .

We note that, even in the cases where the warp factor is non-trivial, W_{harm} can be used to construct the following solutions of the equations of motion:

$$B_2 = \theta(x) W_{\text{harm}} , \quad (34)$$

where $\theta(x)$ is a function of the Minkowski coordinates only. An analogous solution also exists for C_2 . In the equation of motion

$$d \star dB_2 = 0 \quad (35)$$

the warp factor cancels, and it reduces to $d \star_4 d\theta = 0$. As anticipated one can check that this mode of B_2 is not normalizable. This means that it corresponds to a change in the Lagrangian of the dual gauge theory. Indeed choosing a constant θ corresponds to changing $g_1^{-2} - g_2^{-2}$, where g_i are the gauge couplings of the $SU(N) \times SU(N)$ gauge theory [4].

We have not been able to find an analytic solution to (27) and (28) with the warp factor (10). However we can analyze the behavior of the solutions in the asymptotic regimes.

4.3 IR limit: $r \rightarrow 0$

The boundary conditions in the IR are such that W approaches the volume form of the finite 2-sphere. This requires that both f_1 and f_2 vanish at the tip of the cone. In the small r limit, (27) reads

$$\partial_r \left(\frac{hr u^2 \partial_r f_1}{2} \right) + \frac{h}{18u^2 r} (r^4 - 36u^4 f_1) = 0 . \quad (36)$$

In turn, for equation (28) we have

$$\partial_{\theta_2} \left[hr^3 \right] + \partial_r \left[\frac{9u^2}{6} hr^3 \partial_r f_2 \right] = 0 . \quad (37)$$

In order to proceed further, we need the explicit form of the warp factor in the IR. It is known from [9] that its behavior depends crucially on the point in the sphere we are considering. As argued in [9], for a small distance y from the north pole, $h \sim y^{-4}$. For small r and θ_2 , $y^2 \sim 2r^2/3 + u^2 \theta_2^2$, and thus

$$h \sim \frac{9L^4}{4(r^2 + 3u^2 \theta_2^2/2)^2} . \quad (38)$$

In order to analyze the behavior of the fluctuation near the north pole, we will start by setting $\theta_2 = 0$ and keeping a very small r . In this case from (38) $\partial_{\theta_2} h = 0$. Upon taking $h = \frac{9L^4}{4r^4}$, (37) becomes

$$\partial_r \left(\frac{\partial_r f_2}{r} \right) = 0 . \quad (39)$$

The appropriate solution is $f_2 \sim r^2$ which indeed vanishes when r approaches zero. We turn now to equation (36) for $\theta_2 = 0$. It is easy to show that $f_1 = \frac{r^4}{36u^4}$. Note that f_1 also goes to zero with r .

It is instructive to consider an alternative way of reaching the north pole, namely running along the S^2 at $r = 0$ towards $\theta_2 = 0$. From (36) we see that setting $r = 0$ requires us to set $f_1 = 0$ in the whole S^2 . Also at $r = 0$ and small θ_2 , the warp factor in (38) leads to $h = L^4(u\theta_2)^{-4}$. Then equation (37) reduces to

$$\frac{1}{u^2 \theta_2^4} \partial_r (r^3 \partial_r f_2) = 0 . \quad (40)$$

Therefore $\partial_r (r^3 \partial_r f_2) = 0$ at $r = 0$, which sets f_2 to be constant which we choose to vanish.

4.4 The UV limit: $r \rightarrow \infty$

For large r the warp factor approaches that of the singular conifold, namely $h = \frac{L^4}{r^4}$. However in this case the subleading corrections can be written as a series expansion in powers of $1/r$, so it is possible to truncate the series at a certain l . Let us consider not just the pure UV behavior but also the first correction, which already exhibits angular dependence

$$h = \frac{L^4}{r^4} + \frac{9u^2 L^4 \cos \theta_2}{r^6} + \dots . \quad (41)$$

Denoting f_1 and f_2 as the relevant functions in the large r region, it is straightforward to see that they should be of the form

$$f_1 = \frac{1}{2} + F_1(\theta_2, r) , \quad f_2 = F_2(\theta_2, r) ; \quad (42)$$

where F_i are terms which vanish in the asymptotic limit. Therefore both f_1 and f_2 satisfy the required boundary conditions in the UV. The equations (27), (28) read

$$\partial_r \left[\frac{\partial_r F_1}{12r} \right] + \frac{1}{3r^3} \left[-2F_1 + \frac{3}{2} \frac{\partial_{\theta_2} (\sin \theta_2 \partial_{\theta_2} F_1)}{\sin \theta_2} + \frac{\partial_{\theta_2} (F_2 \sin \theta_2)}{\sin \theta_2} \right] - \frac{2u^2}{r^5} = 0 ; \quad (43)$$

and

$$\partial_r \left[\frac{\partial_r F_2}{6r} \right] + \partial_{\theta_2} \left[\left(\frac{1}{r^3} + \frac{(9u^2 \cos \theta_2 - 6u^2)}{r^5} \right) \left(\frac{1}{2} - F_1 + \frac{\partial_{\theta_2} (F_2 \sin \theta_2)}{\sin \theta_2} \right) \right] = 0 . \quad (44)$$

The form of these equations suggests that

$$F_1 = \frac{u^2}{r^2} A(\theta_2) , \quad F_2 = \frac{u^2}{r^2} B(\theta_2) . \quad (45)$$

Then (43) reduces to

$$\frac{\partial_{\theta_2} (\sin \theta_2 \partial_{\theta_2} A)}{\sin \theta_2} + \frac{2}{3} \frac{\partial_{\theta_2} (B \sin \theta_2)}{\sin \theta_2} - 4 = 0 , \quad (46)$$

which can be easily integrated to give the first order equation

$$\frac{1}{2} \sin \theta_2 \partial_{\theta_2} A + \frac{B}{3} \sin \theta_2 = -2 \cos \theta_2 + k , \quad (47)$$

where k is a constant of integration. We can now plug this into (44), and get a single differential equation (which is third order) for A

$$\frac{4k}{\sin \theta_2} - \frac{8 \cos \theta_2}{3 \sin \theta_2} - \frac{9}{2} \sin \theta_2 - \frac{5}{3} \partial_{\theta_2} A - \frac{3}{2} \partial_{\theta_2} \left[\frac{\partial_{\theta_2} (\sin \theta_2 \partial_{\theta_2} A)}{\sin \theta_2} \right] = 0 . \quad (48)$$

Let us mention that it is possible to write down the next order terms, which will still depend on rational powers of r . It suggest that the next correction is going to go like $\frac{1}{r^4}$. However at the next following order, we will already encounter irrational powers of r coming from the fact that h contains them. (48) is most easily solved in terms of $F = \sin \theta_2 \partial_{\theta_2} A$ through the equation

$$\begin{aligned} \sin \theta_2 \partial_{\theta_2} A &= -\tilde{A} {}_2F_1 \left(-\frac{5}{6}, \frac{1}{3}, \frac{1}{2}; \cos^2 \theta_2 \right) \\ &\quad - \tilde{B} {}_2F_1 \left(-\frac{1}{3}, \frac{5}{6}, \frac{3}{2}; \cos^2 \theta_2 \right) \cos \theta_2 + \frac{27}{8} \sin^2 \theta_2 - \frac{8}{5} \cos \theta_2 + \frac{12k}{5} . \end{aligned} \quad (49)$$

where now \tilde{A} and \tilde{B} are more integration constants. Using (48) in (47) gives $B \sin \theta_2$.

For reasons that will be clarified later we would like to keep $\partial_{\theta_2}(B \sin \theta_2)$ finite, which in turn requires that $\partial_{\theta_2} A$ is also finite. Now given (49), in order to for this to hold we have to impose boundary conditions such that the r.h.s vanishes at both 0 and π . This fixes

$$\tilde{A} = \frac{2k}{5\sqrt{\pi}} \Gamma\left(\frac{1}{6}\right) \Gamma\left(\frac{3}{4}\right), \quad \tilde{B} = \frac{24}{5\sqrt{\pi}} \Gamma\left(\frac{2}{3}\right) \Gamma\left(\frac{11}{6}\right). \quad (50)$$

Let us point out that in the asymptotic UV region we find $W = \omega_2$, exactly as in the unwarped toy model.

5 Spontaneous Breaking of the Baryonic Symmetry

In a field theory with spontaneously broken $U(1)$ symmetry, the classical value of the $U(1)$ current is

$$J_\mu^{cl} \sim \frac{i}{2} (\Phi^* \partial_\mu \Phi - \Phi \partial_\mu \Phi^*) = |\Phi_0|^2 \partial_\mu \pi(x), \quad (51)$$

where we substituted $\Phi = \Phi_0 e^{i\pi(x)}$, and $\pi(x)$ is the Goldstone field. Let us show how this expectation value appears for the $U(1)$ baryonic current using the AdS/CFT correspondence. At large r the perturbation δF_5 behaves as

$$\delta F^{(5)} \rightarrow (1 + \star) r^{-3} dr \wedge dp \wedge \omega_3. \quad (52)$$

Therefore the leading term in $\delta C^{(4)}$ contains $r^{-2} dp \wedge \omega_3$ (note that this corresponds to a different gauge choice from that in (23)). We also know that the massless gauge field A_μ^B dual to the baryonic current enters as $\delta C^{(4)} = A^B \wedge \omega_3$.

It follows that

$$A_\mu^B(r) \rightarrow r^{-2} \partial_\mu p. \quad (53)$$

This is a normalizable perturbation near the boundary of AdS₅ that, through the AdS/CFT dictionary [8], implies a relation of the form (51).³

As for any Goldstone boson, it is interesting to determine its “decay constant” f_p which appears in the 4-d effective action as

$$f_p^2 \int d^4x dp \wedge \star_4 dp \quad (54)$$

In deriving this action from dimensional reduction of the type IIB action, we face the usual problem that $\delta F^{(5)}$ containing the Goldstone boson perturbation is self-dual so that $\delta F^{(5)} \wedge \star \delta F^{(5)} = 0$. Instead we will adopt the approach of removing the self-duality constraint and using $\delta F^{(5)} = dp \wedge h \star_6 W$. Now the action no longer vanishes, and we find

$$f_p^2 \sim \frac{1}{g_s^2} \int h W \wedge \star_6 W \quad (55)$$

Note that from (29), $f_p^2 \sim g_s^{-2} I$. Therefore, in the warped case, f_p^2 is a finite quantity determined by the minimum of the functional I . With the analysis of the leading asymptotic corrections we can examine the UV behavior of the integrand in I

$$\int dr \sin \theta_2 \left(\frac{L^4}{6r^3} + \frac{L^4 u^2}{6r^5} \left(9 \cos \theta_2 + 2 \frac{\partial_{\theta_2}(B \sin \theta_2)}{\sin \theta_2} \right) + \mathcal{O}(r^{-7}) \right) \quad (56)$$

³ Essentially the same argument applies to the breaking of the baryonic symmetry in the cascading gauge theory. Using the perturbation containing the Goldstone mode found in [16] we observe the same asymptotic behavior as in (52) up to powers of $\ln r$ characteristic of the cascading theory. This again implies the expectation value of the baryonic current (51).

Note that with the choice of boundary conditions (50), we find

$$\partial_{\theta_2} \left(B \sin \theta_2 \right) |_{\theta=0,\pi} = 0 . \quad (57)$$

Therefore the first correction of order $\mathcal{O}(r^{-5})$ in (56) vanishes upon integration. Note also that (50) also renders $\partial_{\theta_2} A$ finite through the relationship in (49). Indeed with this choice of boundary conditions, the δF_5 remains under control on the whole S^2 .

For scales larger than the resolution length u , we expect the geometry to approach that of the singular conifold. In turn, for scales smaller than u , we expect the resolution of the geometry to take over and smoothly close the cone. Thus, in order to estimate the decay constant, we will take the asymptotic value (56) and cut off the radial integral at $r \sim u$. We find the decay constant goes like

$$f_p^2 \sim \frac{L^4}{u^2 g_s^2} \sim \frac{N}{g_s u^2} \quad (58)$$

Therefore f_p blows up in the limit where u vanishes. Using this we can define a normalized Goldstone boson field $\tilde{p} = p f_p$, in terms of which the VEV of the current takes the canonical form $\langle J_\mu^B \rangle \sim f_p \partial_\mu \tilde{p}$, in agreement with (51).

5.1 Global strings from baryonic symmetry breaking

As argued in section 3, a $D3$ -brane wrapped over the 2-cycle of the resolved conifold is dual to the global string in the gauge theory which arises due to the baryonic symmetry breaking. Now that we have found the supergravity fluctuation (23) of the 4-form gauge potential, which contains the Goldstone boson, we can provide further support for this identification.

The string is charged under the 2-form potential a_2 in Minkowski space. Writing the Minkowski metric as

$$dx_{1,3}^2 = -dt^2 + dx^2 + d\rho^2 + \rho^2 d\theta^2 , \quad (59)$$

we see that a string extended along t, x and localized at $\rho = 0$ produces $a_2 = a(\rho) dt \wedge dx$. The equation of motion following from the 4-dimensional effective action (54) can be recast in terms of a , and reads

$$-\partial_\rho (\rho \partial_\rho a) \sim \frac{T_s}{f_p^2} n \delta(\rho) , \quad (60)$$

where n is the integer winding number of $D3$ around the S^2 . Here T_s stands for the 4d tension of the string in (14). This equation can be integrated to give

$$a(\rho) \sim \frac{T_s}{f_p^2} n \log \rho . \quad (61)$$

We also note that $\star_4 da_2 = d(T_s n \theta / f_p^2)$, giving $p = T_s n \theta / f_p^2$. This exhibits the expected monodromy of the Goldstone boson upon encircling the string. Furthermore we can compute the energy density due to the back-reaction of the string:

$$\mathcal{E} = f_p^2 \int^{\rho_{\max}} (\partial_\mu p)^2 \rho d\rho d\theta \sim \frac{2\pi^2 T_s^2}{f_p^2} n^2 \log \rho_{\max} . \quad (62)$$

Thus we see that the energy density exhibits a logarithmic divergence, which is typical for a global string of codimension two. It is worth noting that \mathcal{E} is of order N^0 , while both f_p^2, T_s are of order N .

5.2 The phase of the baryonic condensate

As shown in [9] the expectation values of baryonic operators can be deduced from the action of certain Euclidean $D3$ -branes ($E3$ -branes). The particular warped resolved conifold studied in [9], corresponding to all $D3$ -branes placed at the north pole of the S^2 at $r = 0$, is dual to the $SU(N) \times SU(N)$ gauge theory with a VEV for a bi-fundamental field B_2 . To calculate the VEV of the gauge invariant baryonic operator $\det B_2$, we need to consider $e^{-S_{E3}} = e^{-S_{DBI}}e^{-S_{CS}}$ for the $E3$ -brane wrapping the 4-chain with internal coordinates $r, \psi, \theta_1, \phi_1$, located at fixed θ_2, ϕ_2 . The phase of $\langle \det B_2 \rangle$ comes from the factor $e^{-S_{CS}}$, since for a Euclidean embedding the Chern-Simons term is imaginary. This fact was important also for computing baryon VEV's in the cascading gauge theory, where they were related to Euclidean $D5$ -branes [17]. For the resolved conifold case, where baryonic operators correspond to $D3$ -branes,

$$S^{CS} = iT_3 \int P[C^{(4)}] . \quad (63)$$

We expect that the phase of the baryonic condensate is proportional to the Goldstone boson field p . Indeed, using the $\delta C^{(4)}$ (23) containing the Goldstone boson, we find that only the second term contributes to the integral:

$$S^{CS} = i16\pi^2 T_3 p \int_0^\infty dr \frac{hr^3}{3(r^2 + 6u^2)} \left(1 - f_1 + \sin^{-1} \theta_2 \partial_{\theta_2} (f_2 \sin \theta_2) \right) . \quad (64)$$

In order to check the convergence of the integral, we note that in the UV f_2 vanishes and f_1 approaches 1/2. The integrand approaches $L^4/(6r^3)$, rendering the integral convergent. We can estimate the integral as before, cutting-off at $r = u$ and using the asymptotic values for f_1, f_2 . We find

$$S^{CS} \sim i \frac{9pN}{8u^2} . \quad (65)$$

Thus although we cannot perform the full integral, we see that the phase of the baryonic condensate behaves as we expected, namely it is proportional to the axion p .

6 Final comments

In this note we studied breaking of the $U(1)$ baryonic symmetry in the warped resolved conifold, and the Goldstone boson it produces. This field might acquire some non-trivial monodromy in the 4-dimensional Minkowski space and generate a global string. We have found such a string and proposed the equations which the axionic Goldstone boson should satisfy. Our results depend on the existence of a very particular mode involving a 2-form in the resolved conifold. This 2-form, W , has to interpolate between the volume form of the finite S^2 at the bottom of the conifold, and the ω_2 2-form in the singular conifold to which the geometry asymptotes. Unfortunately, we have not been able to find this 2-form analytically in the warped case. However, we have provided some evidence that, in the asymptotic regimes, the equations admit a solution which behaves as expected. In any case we have shown that the equations which this 2-form satisfies can be derived from minimizing a functional I . With the appropriate boundary conditions this functional remains finite, supporting the existence of such a 2-form. Additionally, since its norm is given by I , our fluctuation is a normalizable mode. The value of I also controls the decay constant f_p of the Goldstone boson. We note that I is divergent for the unwarped solution and is finite only for the warped solution.

It would be interesting to find the full solution for W . Since we have reduced finding f_1 and f_2 to a variational problem, it should be possible to use a numerical relaxation method with the functional I .

It is worth stressing that the discussion in this note concerns the case of non-compact warped conifold dual to the gauge theory where the baryonic symmetry is global. It would be interesting to consider embedding this set-up into a full string compactification. In that case one would expect the baryonic symmetry to be gauged, and our strings to appear as local ones. The discussion goes in very much the same spirit as that in the case of a pure CY compactification [20, 21]. Since we are considering the resolved conifold, the would-be gauge $U(1)_B$ is in the Higgs phase. In turn this implies a linear potential for the monopole-antimonopole interaction. One would be tempted to identify the string connecting the monopoles with our string, which would then appear as a local string.⁴ However the local string is actually a combination of strings obtained by wrapping branes on a set of cycles that sum up to zero in homology. It would be interesting to adapt the analysis in [20, 21] to the case at hand, in which there is a non-trivial warp factor and 5-form flux. Also, the $\delta C^{(4)}$ mode found here might be responsible for a higher-form mediation of SUSY breaking along the lines of [22]. A key ingredient in this set-up is the existence of two homologous 2-cycles allowing for a combination of two $U(1)$ gauge fields to remain massless. This combination appears in the low energy theory and can potentially mediate SUSY breaking between the hidden and visible sectors. Perhaps such a construction can be implemented in a flux compactification with two different warped resolved conifold throats separated by some distance within the compact dimensions.

A related motivation for our work is that the partially wrapped $D3$ -branes may provide models for cosmic strings in warped compactifications with resolved conifold throats. In such models, the smallness of the string tension is not due to the warp factor, but is rather due to the smallness of the resolution parameter u which also sets the scale of the string tension. Such a construction seems to provide a realization of the unparticle physics scenario [11], since after the RG flow to the IR, the resulting theory contains an interacting CFT (the $\mathcal{N} = 4$ theory) coupled to a non-conformal sector whose scale is set by u .

It would also be interesting to extend the results obtained here to more general resolved CY cones, such as those discussed in [23].

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⁴At the SUGRA level, similar strings to these have been discussed in for example [18, 19].

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